

DOMAIN REDUCTION METHOD FOR
THREE-DIMENSIONAL EARTHQUAKE MODELING IN
LOCALIZED REGIONS. PART I: THEORY

Jacobo Bielak, Kostas Loukakis, Yoshiaki Hisada, and Chiaki Yoshimura

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ABSTRACT

This paper reports on the development of a modular two-step finite element methodology for modeling earthquake ground motion in highly heterogeneous localized regions with large contrasts in wavelengths. We target complex geological structures such as sedimentary basins and ridges that are some distance away from the earthquake source. We overcome the problem of multiple physical scales by subdividing the original problem into two simpler ones. The first is an auxiliary problem that simulates the earthquake source and propagation path effects with a model that encompasses the source and a background structure from which the localized feature has been removed. The second problem models local site effects. Its input is a set of equivalent localized forces derived from the first step. These forces act only within a single layer of elements adjacent to the interface between the exterior region and the geological feature of interest. This enables us to reduce the domain size in the second step. If the background subsurface structure is simple, one can replace the finite element method in the first step with an alternative efficient method. The methodology is illustrated in a companion paper (Yoshimura *et al.*, 2002) for several 3D problems of increasing physical and computational complexity. We consider first a flat-layered, stratigraphic system. For this simple case, the first step can be carried out by means of 3D Green's function evaluations. The extension to more general problems is illustrated by two examples: a basin, and a hill, with the same background stratigraphy. To verify the two-step procedure with a problem for which the finite elements method is used throughout, we model ground motion in a small region of the Los Angeles Basin, using both the two-step domain reduction method, and the traditional approach in which the computational domain

contains both the source and the geological region of interest.

INTRODUCTION

The last ten years have seen tremendous growth in the development of physics-based three-dimensional models for simulating earthquake ground motion in seismic regions. During this period, numerical modeling methods for anelastic wave propagation that take into consideration the earthquake source, propagation path, and local site effects have become increasingly available. There are several types of such methods. Boundary element and discrete wavenumber methods have been popular for moderate-sized problems with relatively simple geometry and geological conditions (e.g., Mossessian and Dravinski, 1987; Kawase and Aki, 1990; Hisada *et al.*, 1993; Sánchez-Sesma and Luzón, 1995; Bouchon and Barker, 1996). Finite differences (e.g., Frankel and Vidale, 1992; Frankel, 1993; Graves, 1993, 1996; Olsen *et al.*, 1995; Pitarka, 1999; Stidham *et al.*, 1999; Sato *et al.*, 1999) and finite elements (e.g., Lysmer and Drake, 1971; Toshinawa and Ohmachi, 1992; Bao, 1998; Bao *et al.*, 1998; Aagaard *et al.*, 2001) are better suited for larger-sized problems that involve realistic basin models with highly heterogeneous materials, due to their flexibility and simplicity. Computer codes based on these methods have been used successfully to model earthquake ground motion in a variety of applications; for instance, near-source ground motion (e.g., Wald and Heaton, 1994), basin structure and directivity effects (e.g., Olsen and Archuleta, 1996; Pitarka *et al.*, 1998); and edge effects (e.g., Kawase, 1996; Hisada *et al.*, 1998; Aagaard *et al.*, 2001).

Despite the recent advances in ground motion simulation capabilities for earthquake excitation (Fig. 1), researchers nowadays are still forced to make restrictive simplifications and approximations in “3D simulations”, such as limiting the maximum frequency or lowest wave velocities that can be considered. One reason is that most methods currently in use for large-size problems are based on uniform structured grids. The grid size, which is proportional to the lowest shear wave velocity in the model and inversely proportional to the highest frequency of interest, is held constant throughout the computational domain even if the softest soils occupy only a small region. Finite elements and other methods with irregular grids (Bao *et al.*, 1998; Pitarka, 1999; Aoi and Fujiwara, 1999; Kristek *et al.*, 1999; Oprsal and Zahradník, 1999, 2002) are more flexible, as they can better tailor the mesh size to the local wavelength of the propagating waves. Regardless of their differences, one feature that is common to traditional finite difference and finite element methods is that the ground motions near the causative fault, and those along the propagation path, and within the region of interest are all calculated simultaneously, using a single model that encompasses the whole geological structure, from the fault region to the region of interest. Thus, source, propagation path, and local site effects are determined all at once. This single-pass approach works well for many applications. However, if the source is far from the region of interest, the problem size becomes exceedingly large and the methods become ineffective.

An alternative formulation, which avoids the need to represent accurately the geometric and material properties of the whole region within a single model, consists in subdividing the problem into two sequential parts. First, one considers a background structure from

which the localized geological features have been deleted, and calculates the corresponding ground motion. This computation requires a grid or mesh that is only as fine as dictated by the softest material in the background model, and needs to be performed only once for a specified earthquake source. In a second step, only a reduced region of interest which contains the localized feature is modeled to the desired accuracy. The ground motion obtained in the first step is used to determine a set of localized equivalent forces, which are then applied as input over a computational domain that is only slightly bigger than the geologic feature. Only the second part of the computation needs to be repeated if any system parameters within the region of interest need to be varied. Such two-step procedure was developed by Bielak and his co-workers in the framework of the finite element method, originally for building-soil-foundation interaction problems (Bielak and Christiano, 1984; Cremonini *et al.*, 1988), and later applied to ground motion modeling of two-dimensional sedimentary valleys in a halfspace due to incident plane SV-waves (Loukakis, 1988; Loukakis and Bielak, 1994a, 1994b). Similar procedures have been presented by Clough and Penzien (1975), Kausel *et al.* (1978), and Aydinoglu (1980, 1993), under some restrictive assumptions. Other researchers have developed alternative two-step or hybrid procedures that make use of combinations of different computational methods; e.g., the wavenumber method and finite differences (Zahradník and Moczo, 1996); modal summation method and finite differences (Regan and Harkrider, 1989; Fäh *et al.*, 1993, 1994); finite element and boundary integral methods (Mita and Luco, 1987; Bielak *et al.*, 1991); and wavenumber method, finite differences, and finite elements introduced to represent irregular geometries (Moczo *et al.*, 1997). A review of various hybrid methods dating back to 1980 can be found in Moczo

et al. (1997). All these methods were concerned with two-dimensional applications. An extension of the finite difference approach to three dimensions has been recently developed by Oprsal and Zahradník (2002).

In this set of two papers we extend the modular two-step procedure developed by Bielak and his co-workers to three-dimensional problems. We use a finite element formulation in which the primary unknowns are the total wave field within the domain that contains the localized structure and a scattered wave field in the exterior domain. This requires that we store the free-field displacements from the background structure only in a single layer of elements at the interface of the two domains in the first step, and enables us to reduce the size of the computational domain in the second step. This methodology, which we call the domain reduction method or DRM, is capable of efficiently modeling three-dimensional wave fields for an arbitrary earthquake source in highly heterogeneous geological systems with large localized impedance contrasts and arbitrary shapes. To validate this procedure, in a companion paper (Yoshimura *et al.*, 2002) we consider a flat-layered system for which a solution can be obtained readily by evaluating the corresponding Green's functions (Hisada, 1994, 1995) and illustrate its applicability to more general problems with two examples, one involving an idealized basin and other a hill. To verify the two-step procedure with a problem for which the finite elements method is used in both steps, we model ground motion in a small region of the Los Angeles Basin, using both the two-step domain reduction method, and the traditional approach in which the computational domain contains both the source and the geological region of interest. While the methodology is applicable to elastic, anelas-

tic, and inelastic problems, only the former are considered explicitly in this set of papers for clarity of presentation. In the concluding remarks, however, we discuss briefly the extension to the more general situations. It will be seen that the use of the DRM can be advantageous even in situations in which the causative fault is not far from the region of interest.

FORMULATION OF DOMAIN REDUCTION METHOD

We treat the problem of a semi-infinite seismic region that contains localized geological features such as sedimentary valleys and ridges as well as seismically active faults, as shown in Fig. 1, under earthquake excitation. The geometry is arbitrary; the material is linearly elastic, and the earthquake excitation is prescribed as a kinematic source along a pre-determined fault.

Since the causative fault can be far from the geological features, we wish to define a new problem in which the excitation is brought closer to the region of interest. Such transfer, of course, needs to be performed in a way that the resulting ground motion within the region of interest is identical to that due to the original source. To fix ideas, suppose the new excitation is to be specified on the fictitious surface Γ shown in Figure 2a. This interface divides the seismic region into two parts: Ω , which contains the geological features of interest, and Ω^+ , the semi-infinite exterior subdomain, which includes the fault. We will determine the appropriate expressions for the equivalent excitation using a finite element formulation. First, the original semi-infinite region needs to be truncated for computational reasons. This is indicated in Fig. 2 by the inclusion of the outer boundary Γ^+ . We assume, for the time being, that it is far enough from the fault that no waves reflected from Γ^+ reach Ω within the time interval under consideration. This assumption will be removed later.

Let the vector field of nodal displacements in the interior domain Ω , the exterior domain Ω^+ , and the boundary between them, Γ , be denoted, respectively, by u_i (interior), u_e (ex-

terior), and u_b (boundary), as shown in Fig. 2a. The seismic excitation is prescribed as a kinematic source, defined by the jump of the tangential displacements across the fault; the normal displacements and tractions remain continuous. This excitation can be equivalently specified by means of body forces (e.g., Aki and Richards, 1980). With this representation, a standard application of Galerkin's ideas with finite element spatial discretization yields a set of nodal forces, P_e , which act in the vicinity of the fault (e.g., Bao,1998).

Rather than analyzing simultaneously the entire domain, which includes both the fault and the localized structure, we would like to focus on the response of a smaller region restricted to a neighborhood of the local structure. To this end, we partition the total domain into two separate subdomains as depicted in Fig. 2b. One contains the fault and the other the localized geological feature. The displacements u_b are continuous across Γ , and P_b are the nodal forces transmitted by Ω^+ onto Ω .

The ground motion within the entire computational domain is governed by Navier's equations of elastodynamics. When these equations are discretized spatially by finite elements in Ω and Ω^+ , they can be expressed in partitioned form as:

$$\begin{bmatrix} M_{ii}^{\Omega} & M_{ib}^{\Omega} \\ M_{bi}^{\Omega} & M_{bb}^{\Omega} \end{bmatrix} \begin{Bmatrix} \ddot{u}_i \\ \ddot{u}_b \end{Bmatrix} + \begin{bmatrix} K_{ii}^{\Omega} & K_{ib}^{\Omega} \\ K_{bi}^{\Omega} & K_{bb}^{\Omega} \end{bmatrix} \begin{Bmatrix} u_i \\ u_b \end{Bmatrix} = \begin{Bmatrix} 0 \\ P_b \end{Bmatrix}, \text{ in } \Omega \quad (1)$$

and

$$\begin{bmatrix} M_{bb}^{\Omega^+} & M_{be}^{\Omega^+} \\ M_{eb}^{\Omega^+} & M_{ee}^{\Omega^+} \end{bmatrix} \begin{Bmatrix} \ddot{u}_b \\ \ddot{u}_e \end{Bmatrix} + \begin{bmatrix} K_{bb}^{\Omega^+} & K_{be}^{\Omega^+} \\ K_{eb}^{\Omega^+} & K_{ee}^{\Omega^+} \end{bmatrix} \begin{Bmatrix} u_b \\ u_e \end{Bmatrix} = \begin{Bmatrix} -P_b \\ P_e \end{Bmatrix}, \text{ in } \Omega^+ \quad (2)$$

In these equations, the matrices M and K denote mass and stiffness matrices, the subscripts i , e , and b refer to nodes in either the interior or the exterior domain or on their common boundary, and the superscripts Ω and Ω^+ refer to the domains over which the various matrices are defined.

The traditional form of the equation of motion for the total domain is obtained by adding (1) and (2):

$$\begin{bmatrix} M_{ii}^{\Omega} & M_{ib}^{\Omega} & 0 \\ M_{bi}^{\Omega} & M_{bb}^{\Omega} + M_{bb}^{\Omega^+} & M_{be}^{\Omega^+} \\ 0 & M_{eb}^{\Omega^+} & M_{ee}^{\Omega^+} \end{bmatrix} \begin{Bmatrix} \ddot{u}_i \\ \ddot{u}_b \\ \ddot{u}_e \end{Bmatrix} + \begin{bmatrix} K_{ii}^{\Omega} & K_{ib}^{\Omega} & 0 \\ K_{bi}^{\Omega} & K_{bb}^{\Omega} + K_{bb}^{\Omega^+} & K_{be}^{\Omega^+} \\ 0 & K_{eb}^{\Omega^+} & K_{ee}^{\Omega^+} \end{bmatrix} \begin{Bmatrix} u_i \\ u_b \\ u_e \end{Bmatrix} = \begin{Bmatrix} 0 \\ 0 \\ P_e \end{Bmatrix} \quad (3)$$

Now, to transfer the seismic excitation from the fault to Γ , we consider an auxiliary problem in which the exterior region and the material therein, as well as the causative fault, are identical to those of the original problem. The interior domain, now denoted as Ω_0 , is, however, a simpler, background structure, that does not include the localized geological features. This is illustrated in Fig. 3a. Ω_0 is chosen such that the new problem defined over the total domain $\Omega_0 \cup \Omega^+$ is easier to solve than the original problem. We denote by u_i^0 , u_b^0 , u_e^0 , and P_b^0 the corresponding nodal displacements and the interface forces, as shown in Fig. 3b. The subscripts i , b , and e , have the same meaning as before. After spatial discretization, the equations of motion in Ω^+ for the auxiliary problem can be written as:

$$\begin{bmatrix} M_{bb}^{\Omega^+} & M_{be}^{\Omega^+} \\ M_{eb}^{\Omega^+} & M_{ee}^{\Omega^+} \end{bmatrix} \begin{Bmatrix} \ddot{u}_b^0 \\ \ddot{u}_e^0 \end{Bmatrix} + \begin{bmatrix} K_{bb}^{\Omega^+} & K_{be}^{\Omega^+} \\ K_{eb}^{\Omega^+} & K_{ee}^{\Omega^+} \end{bmatrix} \begin{Bmatrix} u_b^0 \\ u_e^0 \end{Bmatrix} = \begin{Bmatrix} -P_b^0 \\ P_e \end{Bmatrix} \quad (4)$$

The partitioned mass and stiffness matrices, as well as P_e , are the same as in (2) since the material properties in Ω^+ and the earthquake source are identical in both cases.

From the second equation in (4), we can now express the nodal forces P_e in terms of the free field, as follows:

$$P_e = M_{eb}^{\Omega^+} \ddot{u}_b^0 + M_{ee}^{\Omega^+} \ddot{u}_e^0 + K_{eb}^{\Omega^+} u_b^0 + K_{ee}^{\Omega^+} u_e^0 \quad (5)$$

Then, by substituting (5) into (3), we can solve for the displacements u_i , u_b , and u_e for the complete domain. This formulation by itself, however, offers no advantage over the traditional approach since (5) includes the terms $M_{ee}^{\Omega^+} \ddot{u}_e^0$ and $K_{ee}^{\Omega^+} u_e^0$, which require that the free field u_e^0 be stored throughout the domain Ω^+ . This entails an undue computational effort.

To simplify the analysis, we use a transformation of variables, by which we express the total displacement u_e as the sum of the free field due to the background structure and the residual field due to the localized geological feature:

$$u_e = u_e^0 + w_e \quad (6)$$

That is, the residual field w_e is the relative displacement field with respect to the reference free field u_e^0 .

Then, substituting (6) into (3), and writing the terms that contain the free field on the

right side, results in:

$$\begin{aligned}
& \begin{bmatrix} M_{ii}^{\Omega} & M_{ib}^{\Omega} & 0 \\ M_{bi}^{\Omega} & M_{bb}^{\Omega} + M_{bb}^{\Omega+} & M_{be}^{\Omega+} \\ 0 & M_{eb}^{\Omega+} & M_{ee}^{\Omega+} \end{bmatrix} \begin{bmatrix} \ddot{u}_i \\ \ddot{u}_b \\ \ddot{u}_e \end{bmatrix} + \begin{bmatrix} K_{ii}^{\Omega} & K_{ib}^{\Omega} & 0 \\ K_{bi}^{\Omega} & K_{bb}^{\Omega} + K_{bb}^{\Omega+} & K_{be}^{\Omega+} \\ 0 & K_{eb}^{\Omega+} & K_{ee}^{\Omega+} \end{bmatrix} \begin{bmatrix} u_i \\ u_b \\ w_e \end{bmatrix} = \\
& \begin{bmatrix} 0 \\ -M_{be}^{\Omega+} \ddot{u}_e^0 - K_{be}^{\Omega+} u_e^0 \\ P_e - M_{ee}^{\Omega+} \ddot{u}_e^0 - K_{ee}^{\Omega+} u_e^0 \end{bmatrix} \quad (7)
\end{aligned}$$

Finally, after substituting for P_e from (5) into (7), we obtain the desired equation:

$$\begin{aligned}
& \begin{bmatrix} M_{ii}^{\Omega} & M_{ib}^{\Omega} & 0 \\ M_{bi}^{\Omega} & M_{bb}^{\Omega} + M_{bb}^{\Omega+} & M_{be}^{\Omega+} \\ 0 & M_{eb}^{\Omega+} & M_{ee}^{\Omega+} \end{bmatrix} \begin{bmatrix} \ddot{u}_i \\ \ddot{u}_b \\ \ddot{u}_e \end{bmatrix} + \begin{bmatrix} K_{ii}^{\Omega} & K_{ib}^{\Omega} & 0 \\ K_{bi}^{\Omega} & K_{bb}^{\Omega} + K_{bb}^{\Omega+} & K_{be}^{\Omega+} \\ 0 & K_{eb}^{\Omega+} & K_{ee}^{\Omega+} \end{bmatrix} \begin{bmatrix} u_i \\ u_b \\ w_e \end{bmatrix} = \\
& \begin{bmatrix} 0 \\ -M_{be}^{\Omega+} \ddot{u}_e^0 - K_{be}^{\Omega+} u_e^0 \\ M_{eb}^{\Omega+} \ddot{u}_b^0 + K_{eb}^{\Omega+} u_b^0 \end{bmatrix} \quad (8)
\end{aligned}$$

The mass matrix and stiffness matrix in the left hand side of (8) are identical to those of (3). However, the seismic forces P_e on the fault have been replaced by the effective nodal forces P^{eff} , given by:

$$P^{\text{eff}} = \begin{bmatrix} P_i^{\text{eff}} \\ P_b^{\text{eff}} \\ P_e^{\text{eff}} \end{bmatrix} = \begin{bmatrix} 0 \\ -M_{be}^{\Omega+} \ddot{u}_e^0 - K_{be}^{\Omega+} u_e^0 \\ M_{eb}^{\Omega+} \ddot{u}_b^0 + K_{eb}^{\Omega+} u_b^0 \end{bmatrix} \quad (9)$$

These forces have the key property that they involve only the submatrices M_{be} , K_{be} , M_{eb} , and K_{eb} , which vanish everywhere except in a single layer of finite elements in Ω^+ adjacent

to Γ . This small domain lies between Γ and its adjacent surface Γ_e , as shown in Fig. 4. Therefore, the forces P^{eff} act exclusively within that layer. Also, the only wavefield needed to determine P^{eff} is that obtained from the auxiliary problem at the nodes that lie on Γ , Γ_e , and between these surfaces. This localization of the equivalent seismic forces around the geologic feature is the key advantage of the transformation (6).

Another important consequence of (9) is that all the waves in the exterior region Ω^+ will be outgoing. This suggests that for solving (8), the size of the region Ω^+ can be drastically reduced if one is interested only in the ground motion near the localized features, provided suitable absorbing boundaries are used to limit the occurrence of spurious waves. Due to this attractive feature, we name our method Domain Reduction Method (DRM). To emphasize this reduction in size, we will denote the reduced exterior region by $\hat{\Omega}^+$ and its corresponding outer boundary by $\hat{\Gamma}^+$. These results were derived originally in the context of a halfspace and plane wave excitation in a slightly different form (Bielak and Christiano, 1984; Loukakis, 1988; Loukakis and Bielak, 1994). The present derivation is more rigorous and concise, and incorporates explicitly the effect of an extended source on a finite fault. A similar procedure to that in Loukakis (1988) was developed subsequently by Aydinöglu (1993), in the context of soil-structure interaction without explicit treatment of the earthquake source. Instead of using a finite element formulation throughout as in Loukakis (1988) and Loukakis and Bielak (1994), Aydinöglu (1993) used a boundary integral representation for the tractions at the interface between the interior and exterior domains; to make the equations local at the interface, the traction was approximated in the form

of a mass-dashpot-spring, and the material outside the interface Γ was excluded from the computations. It is not clear that this simplification will lead to acceptable approximations. In addition, contrary to our formulation, in which the effective forces depend only on the properties of the material within the region exterior to Ω , the effective forces are defined on a layer within the interior region. This could pose some difficulty if one is interested in considering interior regions that behave nonlinearly, as discussed in the next section. Similar results were derived also by Zahradník and Moczo (1996) for the finite difference method in two dimensions using a rectangular excitation box, and an algorithm similar to that of Alterman and Karal (1968) to do the coupling.

DISCUSSION AND CONCLUDING REMARKS

The results described in the previous paragraphs can be summarized as a two-step procedure for analyzing the earthquake response of localized geological features, as follows. In the first step, I, as shown in Fig. 5a, one starts with a background geological model defined over the domains Ω^+ and Ω_0 , which include the original earthquake source, and defines the boundary Γ of what in step II will be the region of interest. Then one calculates the free-field ground motion u_b^0 and u_e^0 and stores it at all the nodes on the adjacent surfaces Γ and Γ_e as well as at any interior nodes within the finite element layer that lies between them. Suitable absorbing boundary conditions on Γ^+ , such as those proposed by Clayton and Engquist (1977) and Stacey (1988), must be used to keep spurious wave reflections within acceptable limits. Spurious reflections from the absorbing boundaries are generally unavoidable in practice, and can lead to inaccuracies in the numerical results.

If desired, any appropriate method can be used in step I to determine these wavefields in place of the finite element method. The second step is performed on a reduced region $\Omega \cup \hat{\Omega}^+$, which contains the geological features of interest, but not the causative fault, as shown in Fig. 5b. The effective seismic forces P^{eff} are evaluated first, from (9), using u_b^0 and u_e^0 as input. Once these forces have been established, the total wavefields u_i and u_b , and the residual wavefield w_e , defined respectively over Ω , Γ , and $\hat{\Omega}^+$, are obtained by solving (8). Actually, this equation must be modified slightly to incorporate the absorbing boundary conditions on $\hat{\Gamma}^+$. The complete solution u_e within the domain $\hat{\Omega}^+$ can be evaluated simply as $u_e = w_e + u_e^0$.

It is important to emphasize that the DRM is exact to within finite element spatial and time discretization errors. In our implementation we have used piecewise linear finite elements in space, and second-order central differences in time. Thus, our results are second-order accurate in space and time. The main computational efficiency is gained from the use of the finite element method, which allows one to tailor the mesh size to the local wavelength of the propagating wave, and to the reduction of the overall computational domain in the second step, which might be quite dramatic if the source is far, the material within the region of interest is nonlinear, or repeated solutions with minor changes in the parameters are required, as in inverse problems. With the goal of increasing the computational efficiency of 3D simulations, other authors have developed powerful memory optimization methods, especially in the context of the finite difference method, in order to be able to deal with more realistic

material properties and frequencies closer to those of engineering interest. In particular, Moczo *et al.* (1999) use a lossy compression scheme based on a discrete wavelet transform to optimize RAM and disk storage requirements. Combining the compression scheme with the memory optimization of Graves (1996) and the memory variable economization of Day (1998) and Day and Bailey (2001), Moczo *et al.* (2000, 2001) in some cases achieve a total disk plus RAM storage reduction approaching or exceeding an order of magnitude. Such optimization techniques could be applied also to the DRM to attain further storage reductions.

It should also be emphasized that the strict validity of our two-step methodology hinges on the linearity of the material outside the region of interest Ω and on the requirement that the material properties in the exterior region be the same for the auxiliary problem as for the original problem. The material within Ω , on the other hand, can be arbitrary. Notice that its properties do not enter into the computation of P^{eff} . In fact, the two-step procedure we just described would remain valid even if the material in Ω were nonlinear. All that would change in the formulation are the stiffness terms in (3) and (7) pertaining to the interior domain, which would then depend on the solution. More generally, the second term in (1) and in subsequent equations would be a nonlinear function of u_i and u_b , which would depend on the material constitutive equations. Another assumption in our derivation is that the material is non-dissipative. Clearly, adding linear dissipation would not change the essence of the procedure, but merely modify the particular form of the equations. For instance, the addition of viscous damping would result in added terms proportional to velocity in the equations (Loukakis, 1988; Loukakis and Bielak, 1994). More generally, linear viscoelastic

behavior would, rigorously, introduce convolutions into the semi-discretized formulation. Such convolutions, however, can be avoided if one approximates the viscoelastic dependence by a linear combination of multiple relaxation mechanisms that can be expressed in differential form with the aid of auxiliary memory variables (e.g., Day and Minster, 1984; Emmerich and Korn, 1987; Carcione et al, 1988, Moczo *et al.*, 1997). One important drawback encountered originally with these methods was the added computational and storage cost associated with the auxiliary memory variables. This difficulty has been partially removed by the coarse-graining methodology developed by Day (1998), in which only one individual relation mechanism per node produces highly accurate results.

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Author affiliations

Department of Civil and Environmental Engineering

Carnegie Mellon University

Pittsburgh, PA 15213-3890

jbielak@cmu.edu

(J. B.)

Technical Company of General Construction

Isiodou 22, Athens 10674, Greece

klukakis@compulink.gr

(K. L.)

Department of Architecture

Kogakuin University

1-24-2 Nishi-shinjuku, Shinjuku-ku

Tokyo 163-8677, Japan

hisada@cc.kogakuin.ac.jp

(Y. H.)

Technology Center

Taisei Corporation

344-1 Nase-cho, Totsuka-ku

Yokohama 245-0051, Japan

yosimura@eng.taisei.co.jp

(C. Y.)

Figure legends

Figure 1. Schematic of semi-infinite seismic region. It includes causative fault, geological structure and local features.

Figure 2. Truncated seismic region. (a) Outer boundary Γ^+ restricts computations to a finite domain; fictitious interface Γ divides region into two subdomains: Ω^+ , which includes the seismic source, represented by nodal forces P_e , and Ω which contains the localized geological features. (b) Regions partitioned explicitly into two substructures across interface Γ ; P_b are nodal forces transmitted from Ω^+ onto Ω ; $-P_b$ are corresponding reactions from Ω onto Ω^+ ; nodal displacements u_b are required to be continuous across Γ .

Figure 3. Auxiliary seismic region. Localized geological features of actual problem in Ω have been replaced by a simpler background structure over domain Ω_0 . (a) Entire region; (b) Region partitioned into two substructures.

Figure 4. Seismic region with two neighboring surfaces Γ and Γ_e on which effective nodal forces P^{eff} defined by Eq. (9) are to be applied. These forces are equivalent to and replace the original seismic forces P_e , which act in the vicinity of the causative fault.

Figure 5. Summary of two-step domain reduction method (DRM). (a) Step I defines the auxiliary problem over background geological model. Resulting nodal displacements within Γ , Γ_e and the region between them are used to evaluate effective seismic forces P^{eff} required for Step II. (b) Step II, defined over reduced region made up of Ω and $\hat{\Omega}^+$ (a truncated portion of Ω^+). The effective seismic forces P^{eff} are applied within

Γ and Γ_e . The unknowns are: the total displacement fields u_i in Ω and u_b on Γ , and the residual displacements w_e in $\hat{\Omega}^+$.