

Ben Sauerwine  
Quantum Mechanics 3 Homework 7

1) Obtain  $\langle n|U(0,-\infty)|n\rangle$  in second order in  $H_I$ . Extract the terms proportional to  $\frac{1}{\varepsilon^2}$  and  $\frac{1}{\varepsilon}$  in the limit  $\varepsilon \rightarrow 0$ . Compare your result with an expression of the form  $e^{-\frac{i}{\hbar}\Delta\varepsilon_n T} e^{-\frac{\Gamma_n}{2} T}$  with  $T = \frac{1}{\varepsilon}$  and  $\Delta\varepsilon_n$  and  $\Gamma_n$  the energy shift and decay rate, respectively.

Recall that

$$\Gamma_n = \frac{2\pi}{\hbar} \sum_{m \neq n} |\langle m|H_I|n\rangle|^2 \delta(E_m - E_n)$$

and

$$\Delta\varepsilon_n = \langle n|H_I|n\rangle + \lim_{\eta \rightarrow 0} \sum_{m \neq n} \frac{|\langle n|H_I|m\rangle|^2}{E_n - E_m + i\eta}$$

The perturbed state is given by:

$$|\tilde{n}\rangle = U(0,-\infty)|n\rangle = |n\rangle - \frac{i}{\hbar} \int_{-\infty}^0 dt H_I(t)|n\rangle - \frac{1}{2\hbar^2} \int_{-\infty}^0 \int_{-\infty}^0 dt' dt'' (H_I(t')H_I(t'')\theta(t'-t'') + H_I(t'')H_I(t')\theta(t''-t'))|n\rangle + \dots$$

Strategically inserting  $I = \sum_m |m\rangle\langle m|$  and applying  $\langle n|$  I may obtain in the second order:

$$\begin{aligned} \langle n|\tilde{n}\rangle &= 1 - \frac{i}{\hbar} \int_{-\infty}^0 dt' \langle n|H_I(t')|n\rangle \\ &+ \frac{1}{2\hbar^2} \sum_m \int_{-\infty}^0 \int_{-\infty}^0 dt' dt'' (\langle n|H_I(t')|m\rangle\langle m|H_I(t'')\theta(t'-t'') + \langle n|H_I(t'')|m\rangle\langle m|H_I(t')\theta(t''-t'))|n\rangle \end{aligned}$$

Now recalling the form of my interaction Hamiltonian,

$$\langle m|H_I(t)|n\rangle = \langle m|H_I|n\rangle e^{\frac{i}{\hbar}(E_m - E_n)t} e^{\varepsilon t}$$

Above,  $\langle m|H_I|n\rangle$  is now time-independent.

Then I have:

$$\langle n|\tilde{n}\rangle = 1 - \frac{i}{\hbar} \langle n|H_I|n\rangle \int_{-\infty}^0 dt' e^{\varepsilon t'} \\ - \sum_m \frac{|\langle n|H_I|m\rangle|^2}{\hbar^2} \int_{-\infty}^0 \int_{-\infty}^0 dt' dt'' \left( e^{\frac{i}{\hbar}(E_n - E_m)(t' - t'')} e^{\varepsilon t'} e^{\varepsilon t''} \theta(t' - t'') + e^{-\frac{i}{\hbar}(E_n - E_m)(t' - t'')} e^{\varepsilon t'} e^{\varepsilon t''} \theta(t'' - t') \right)$$

Integrating, I see that in this case, the two terms in the second-order integral give the same result.

$$\langle n|\tilde{n}\rangle = 1 - \frac{i}{\hbar} \frac{\langle n|H_I|n\rangle}{\varepsilon} e^{\varepsilon t} - \sum_m \frac{|\langle n|H_I|m\rangle|^2}{\hbar^2} \int_{-\infty}^0 dt \left( \frac{e^{2t\varepsilon} \theta(-t) + e^{\frac{i}{\hbar}(E_n - E_m)t} e^{t\varepsilon} \theta(t)}{-\frac{i}{\hbar}(E_n - E_m) + \varepsilon} \right)$$

Integrating what remains, I have:

$$\langle n|\tilde{n}\rangle = 1 - \frac{i \langle n|H_I|n\rangle}{\hbar \varepsilon} e^{\varepsilon t} - \frac{1}{2\hbar^2} \sum_m \frac{|\langle n|H_I|m\rangle|^2}{\varepsilon \left( -\frac{i}{\hbar}(E_n - E_m) + \varepsilon \right)}$$

Now all that's left to do is take the limit as  $\varepsilon \rightarrow 0$ , and separate the sum over  $m$  into a separate sum for degenerate, identity, and non-degenerate cases (note that except for the  $n = m$  case, there are 2 of each term):

$$\langle n|\tilde{n}\rangle = 1 - \frac{i}{\hbar \varepsilon} \langle n|H_I|n\rangle - \frac{i}{\hbar \varepsilon} \sum_{m \neq n, E_n \neq E_m} \frac{|\langle n|H_I|m\rangle|^2}{(E_n - E_m)} - \frac{1}{2\hbar^2 \varepsilon^2} |\langle n|H_I|n\rangle|^2 - \frac{\pi}{\hbar \varepsilon^2} \sum_{m \neq n, E_n = E_m} |\langle n|H_I|m\rangle|^2$$

Identifying terms, then, I have:

$$\langle n|\tilde{n}\rangle = 1 - \frac{i}{\hbar} T \Delta E_n^{(1)} - \frac{i}{\hbar} T \Delta E_n^{(2)} + \frac{1}{2} \left( -\frac{i}{\hbar} \right)^2 T^2 (\Delta E_n^{(1)})^2 - \frac{\Gamma_n}{2}$$

So, then, I see that these represent several terms from  $e^{-\frac{i}{\hbar} \Delta \varepsilon_n T} e^{-\frac{\Gamma_n T}{2}}$ ; namely, if I let:

$$\Delta \varepsilon_n = \Delta E_n^{(1)} + \Delta E_n^{(2)} + \dots$$

Then the procedure that I followed gives me all terms of the product of these expansions out to terms square in  $H_I$ :

$$e^{-\frac{i}{\hbar}\Delta\varepsilon_n T} e^{-\frac{\Gamma_n}{2} T} = \left( 1 + \left( -\frac{i}{\hbar} \Delta\varepsilon_n \right) T + \frac{1}{2} \left( -\frac{i}{\hbar} \Delta\varepsilon_n \right)^2 T^2 + \dots \right) \left( 1 - \frac{\Gamma_n}{2} T + \dots \right)$$

2) Consider hydrogenic atoms in a cavity in which a photon bath is in equilibrium at temperature  $T$  and consider transitions only between ground and first excited states (only two levels:  $A, B$ ).

a) Follow the steps in class and obtain  $\Gamma_{abs}$  and  $\Gamma_{em}$ . Leave the angular integral

$$\sum_{\lambda} \int \frac{d\Omega}{4\pi} \vec{e}_{\lambda} \cdot \vec{x}_{AB} \text{ unspecified.}$$

$$H_I^{abs} = \frac{1}{\sqrt{V}} \left( -\frac{e}{mc} \right) c_B^+ c_A a_{k,\lambda} \int d^3x U_B^*(\vec{x}) \vec{A}(\vec{x}) \cdot \hat{p} U_A(\vec{x})$$

$$H_I^{em} = \frac{1}{\sqrt{V}} \left( -\frac{e}{mc} \right) c_B c_A^+ a_{k,\lambda}^+ \int d^3x U_A^*(\vec{x}) \vec{A}(\vec{x}) \cdot \hat{p} U_B(\vec{x})$$

$$\hat{p} = -i\hbar \vec{\nabla}_x$$

$$\vec{A}(\vec{x}) = \sum_{k,\lambda} \sqrt{\frac{\hbar c^2}{2\omega(\vec{k})}} \vec{e}_{\lambda}(\vec{k}) e^{i\vec{k} \cdot \vec{x}} \rightarrow \text{dipole approximation} \rightarrow \sum_{k,\lambda} \sqrt{\frac{\hbar c^2}{2\omega(\vec{k})}} \vec{e}_{\lambda}(\vec{k})$$

$$\text{Further, Fermi's Golden Rule says that } \Gamma_x = \frac{2\pi}{\hbar} \sum_f \left| \langle F | H_I^x | I \rangle \right|^2 \delta(\varepsilon_f - \varepsilon_i)$$

I will do absorption first: I seek to use Fermi's Golden Rule to obtain a lifetime.

$$|F\rangle = |B\rangle |n_{k,\lambda} - 1\rangle \quad |I\rangle = |A\rangle |n_{k,\lambda}\rangle$$

Now I see that

$$\langle F | H_I^{abs} | I \rangle = \frac{1}{\sqrt{V}} \langle B | c_B^+ c_A | A \rangle \langle n_{k,\lambda} - 1 | a_{k,\lambda} | n_{k,\lambda} \rangle \left( -\frac{e}{mc} \right) \sqrt{\frac{\hbar c^2}{2\omega(\vec{k})}} \int d^3x U_B^*(\vec{x}) \vec{e}_{\lambda}(\vec{k}) \cdot \hat{p} U_A(\vec{x})$$

However, I notice that

$$\int d^3x U_B^*(\vec{x}) \hat{p} U_A(\vec{x}) = -\frac{im}{\hbar} \int d^3x U_B^*(\vec{x}) [\vec{x}, H_0^m] U_A(\vec{x}) = \frac{im}{\hbar} \langle B | \hat{x} | A \rangle (E_B - E_A)$$

$$\langle n_{k,\lambda} - 1 | a_{k,\lambda} | n_{k,\lambda} \rangle = \sqrt{n_{k,\lambda}}$$

$$\langle B | c_B^+ c_A | A \rangle = 1$$

and so:

$$\langle F | H_I^{abs} | I \rangle = \frac{1}{\sqrt{V}} \frac{im}{\hbar} (E_B - E_A) \sqrt{n_{k,\lambda}} \left( -\frac{e}{mc} \right) \sqrt{\frac{\hbar c^2}{2\omega(\vec{k})}} \vec{e}_{\lambda}(\vec{k}) \cdot \langle B | \hat{x} | A \rangle$$

Then, using  $\hbar\omega(k) = \hbar ck = E_B - E_A$ , I see that:

$$\left| \langle F | H_I^{abs} | I \rangle \right|^2 = \frac{e^2}{2V} (E_B - E_A) n_{k,\lambda} \left| \bar{e}_\lambda(\vec{k}) \cdot \langle B | \hat{x} | A \rangle \right|^2$$

Now in Fermi's Golden Rule,

$$\Gamma_{abs} = \frac{2\pi}{\hbar} \sum_f \frac{e^2}{2V} (E_B - E_A) n_{k,\lambda} \left| \bar{e}_\lambda(\vec{k}) \cdot \langle B | \hat{x} | A \rangle \right|^2 \delta(E_B - E_A - \hbar ck)$$

$$\sum_f \rightarrow \sum_{k,\lambda} \quad \frac{1}{V} \sum_{k,\lambda} \rightarrow \sum_\lambda \int \frac{d^3k}{(2\pi)^3} = \sum_\lambda \int_0^\infty k^2 dk \int \frac{d\Omega}{8\pi^3}$$

$$\Gamma_{abs} = \frac{\pi e^2}{\hbar} (E_B - E_A) n_{k,\lambda} \sum_\lambda \int_0^\infty k^2 \delta(E_B - E_A - \hbar ck) dk \int \frac{d\Omega}{8\pi^3} \left| \bar{e}_\lambda(\vec{k}) \cdot \langle B | \hat{x} | A \rangle \right|^2$$

$$\int_0^\infty k^2 \delta(E_B - E_A - \hbar ck) dk = \frac{(E_B - E_A)^2}{(\hbar c)^3}$$

$$\Gamma_{abs} = \frac{e^2}{2\hbar} \frac{(E_B - E_A)^3}{(\hbar c)^3} n_{k,\lambda} \sum_\lambda \int \frac{d\Omega}{4\pi^2} \left| \bar{e}_\lambda(\vec{k}) \cdot \langle B | \hat{x} | A \rangle \right|^2$$

where now  $\vec{k} = \frac{(E_B - E_A)}{\hbar c} \hat{n}$ .

Emission is very similar; summarizing, the only real differences are:

$$\langle F | H_I^{em} | I \rangle = \frac{1}{\sqrt{V}} \langle A | c_A^\dagger c_B | B \rangle \langle n_{k,\lambda} + 1 | a^{+}_{k,\lambda} | n_{k,\lambda} \rangle \left( -\frac{e}{mc} \right) \sqrt{\frac{\hbar c^2}{2\omega(\vec{k})}} \int d^3x U_A^*(\vec{x}) \bar{e}_\lambda(\vec{k}) \cdot \hat{p} U_B(\vec{x})$$

$$\langle n_{k,\lambda} + 1 | a^{+}_{k,\lambda} | n_{k,\lambda} \rangle = \sqrt{n_{k,\lambda} + 1}$$

Propagating these through (square-roots and absolute values ensure that that the sign remains the same:

$$\Gamma_{em} = \frac{e^2}{2\hbar} \frac{(E_B - E_A)^3}{(\hbar c)^3} (1 + n_{k,\lambda}) \sum_\lambda \int \frac{d\Omega}{4\pi^2} \left| \bar{e}_\lambda(\vec{k}) \cdot \langle B | \hat{x} | A \rangle \right|^2$$

**b) Approximate the angular integral by the Bohr radius of a hydrogenic atom and estimate  $\Gamma_{abs}$ ,  $\Gamma_{em}$  for 300K and 3000K.**

For emission, I see that *within factors that I had to add to make units work*, stimulated emission was very unlikely.

$$\int \frac{d\Omega}{4\pi^2} |\vec{e}_\lambda(\vec{k}) \cdot \langle B | \dot{\vec{x}} | A \rangle|^2 \rightarrow a_0^2$$

$$\Gamma_{em} \rightarrow (1 + n_{k,\lambda}) \frac{e^2}{\hbar c} \left( \frac{\Delta E}{\hbar} \right)^3 \left( \frac{a_0}{c} \right)^2 = (1 + n_{k,\lambda}) \frac{e^2}{\hbar c} \left( \frac{\Delta E}{\hbar} \right)^3 \left( \frac{a_0}{c} \right)^2$$

$$\alpha = \frac{e^2}{\hbar c} = \frac{1}{137}$$

$$\Delta E = \frac{3}{4} (13.6) eV$$

$$a_0 = 52.9 \cdot 10^{-3} nm$$

$$\hbar c = 197.3 eV \cdot nm$$

$$\hbar = 6.582 \cdot 10^{-16} eV \cdot s$$

$$\Gamma_{em} \rightarrow (1 + n_{k,\lambda}) \frac{e^2}{\hbar c} \left( \frac{\Delta E}{\hbar} \right)^3 \left( \frac{a_0}{c} \right)^2 = (1 + n_{k,\lambda}) [8.46 \cdot 10^8 s^{-1}]$$

Using the value for number of particles from the Canonical ensemble:

$$k_B = 8.67 \cdot 10^{-5} \frac{eV}{K}$$

$$n_{k,\lambda} = \frac{e^{-\frac{117647}{T}}}{e^{-\frac{117647}{T}} + 1}$$

$$\Gamma_{em}(300K) \approx 1.69 \cdot 10^9 s^{-1}$$

$$\Gamma_{em}(3000K) \approx 1.69 \cdot 10^9 s^{-1}$$

So I see that the emissive lifetime is on the nanosecond scale, across temperatures. The absorption formula is practically identical, except for the factor arising from spontaneous emission:

$$\Gamma_{abs}(300K) \approx 8.46 \cdot 10^8 s^{-1}$$

$$\Gamma_{abs}(3000K) \approx 8.46 \cdot 10^8 s^{-1}$$

So I see that the expected lifetime of the non-excited state is also very low: the particles rapidly exchange energy.

c) In the same approximation, estimate the radiative lifetime  $\tau = \frac{1}{\Gamma_{em}}$  at  $T = 0$

for the excited state.

In this approximation,  $n_{k,\lambda} \rightarrow 0$  and so

$$\Gamma_{em} \rightarrow (1 + n_{k,\lambda}) \frac{e^2}{\hbar c} \left( \frac{\Delta E}{\hbar} \right)^3 \left( \frac{a_0}{c} \right)^2 = (1 + n_{k,\lambda}) [8.46 \cdot 10^8 \text{ s}^{-1}]$$

$$\tau \rightarrow 1.18 \cdot 10^{-9} \text{ s}$$

or about a nanosecond.

**3) Consider  $N$  atoms with only two levels A and B in a cavity with blackbody radiation in equilibrium at temperature  $T$ .**

**a) If at time  $t = 0$  all the atoms are prepared to be in the level A, give the populations  $N_A(t)$ ,  $N_B(t)$  and show that they approach their equilibrium value exponentially.**

The atoms in the cavity will change population through gain and loss: namely,

$$\frac{dN_A}{dt} = N_B \Gamma_{em} - N_A \Gamma_{abs} \quad \frac{dN_B}{dt} = N_A \Gamma_{abs} - N_B \Gamma_{em}$$

However, since any atom must be in state A or B,

$$\frac{d(N_A + N_B)}{dt} = 0$$

then in equilibrium:

$$\left. \frac{dN_A}{dt} \right|_{eq} = 0$$

or

$$\left. \frac{dN_A}{dt} \right|_{eq} = 0 = N_{B,eq} \Gamma_{em} - N_{A,eq} \Gamma_{abs}$$

$$\frac{N_{B,eq}}{N_{A,eq}} = \frac{\Gamma_{abs}}{\Gamma_{em}}$$

$$\frac{dN_A}{dt} = (N - N_A) \Gamma_{em}$$

This implies that the function  $N_A$  must approach its equilibrium exponentially. Clearly, a similar proof can be given for  $N_B$ .

**b) What are the equilibrium values  $N_{A,eq} = N_A(\infty)$ ,  $N_{B,eq} = N_B(\infty)$  in terms of  $N$ ,  $\Gamma_{abs}$ ,  $\Gamma_{em}$  ?**

There are certainly always  $N$  particles. From part (a), I found that

$$\frac{N_{B,eq}}{N_{A,eq}} = \frac{\Gamma_{abs}}{\Gamma_{em}}$$

Then,

$$\frac{N_{A,eq}}{N - N_{A,eq}} = \frac{\Gamma_{em}}{\Gamma_{abs}}$$

$$N_{A,eq} = N \frac{\frac{\Gamma_{em}}{\Gamma_{abs}}}{1 + \frac{\Gamma_{em}}{\Gamma_{abs}}} \quad N_{B,eq} = N \frac{1}{1 + \frac{\Gamma_{em}}{\Gamma_{abs}}}$$

**c) Consider these to be hydrogenic atoms with only ground and first excited states A, B and assume that the angular average  $\sum_{\lambda} \int \frac{d\Omega}{4\pi} \bar{e}_{\lambda} \cdot \bar{x}_{AB}$  is approximately the Bohr radius. Estimate the equilibrium time scale  $\tau_{eq}$  for  $T = 300K, 3000K$ , where  $\tau_{eq}$  is defined from the exponential approach to the equilibrium of the populations  $\propto e^{-\frac{t}{\tau_{eq}}}$ .  $\tau_{eq}$  is solely a function of  $\Gamma_{em}, \Gamma_{abs}$ .**

Using

$$\frac{dN_A}{dt} = (N - N_A)\Gamma_{em} - N_A\Gamma_{abs}$$

$$\frac{dN_A}{dt} = N\Gamma_{em} - (\Gamma_{em} + \Gamma_{abs})N_A - N_A\Gamma_{abs}$$

Solving this under the initial condition  $N_A(0) = N$ , I have:

$$N_A = \frac{\Gamma_{abs}e^{(-\Gamma_{abs}-\Gamma_{em})t} + \Gamma_{em}}{\Gamma_{abs} + \Gamma_{em}} N$$

The equilibrium timescale should then be like  $\frac{1}{\Gamma_{abs} + \Gamma_{em}}$ : Taking the results from 2b, I have:

$$\frac{1}{\Gamma_{abs} + \Gamma_{em}} = \frac{1}{1.69 \cdot 10^9 + 8.46 \cdot 10^8} s = 3.94 \cdot 10^{-10} s$$

This is practically independent of temperature, since the lifetimes of these states were identical within the proper number of significant figures.