

Problem Source: CMU August 2003 Qualifying Exam

Consider a hollow body at temperature T containing electromagnetic radiation in equilibrium with its walls. The energy spectrum of this so-called blackbody radiation can be measured by observing it through a negligibly small hole in the cavity wall. As argued by Planck, the energy levels of the radiation are quantized and are the same as those of a simple harmonic oscillator of angular frequency ω .

- (a) What are the quantized energy levels of a radiation mode of angular frequency ω ? Use the Boltzmann distribution to find the thermal average number of photons for such a mode.

$$E_n = \hbar\omega\left(n + \frac{1}{2}\right)$$

(ignore zero point)

$$E_n = \hbar\omega n$$

$$Z = \sum_n e^{-\beta E_n} = \sum_n e^{-\beta\hbar\omega n} = \sum_n (e^{-\beta\hbar\omega})^n = \frac{1}{1 - e^{-\beta\hbar\omega}}$$

$$\langle E \rangle = -\frac{\partial}{\partial \beta} \ln Z = \frac{\hbar\omega e^{-\beta\hbar\omega}}{1 - e^{-\beta\hbar\omega}}$$

$$\langle E \rangle = \hbar\omega \langle N \rangle$$

$$\langle N \rangle = \frac{1}{e^{\beta\hbar\omega} - 1}$$

- (b) The density of modes in \vec{k} space is constant. Thus for a cavity of volume V , the number of modes in a volume d^3k near a wave number \vec{k} is just

$$2 \frac{V}{(2\pi)^3} d^3k$$

If c is the speed of light, how many modes are there having angular frequency between ω and $\omega + d\omega$?

In terms of wavelength and frequency, $\omega = ck$. Then, converting the volume-wise density into an integral only over the surface, I have:

$$\rightarrow 2 \frac{V}{(2\pi)^3} (4\pi k^2) dk$$

$$k = \frac{\omega}{c} \quad dk = \frac{1}{c} d\omega$$

for

$$\rightarrow 2 \frac{V}{(2\pi)^3} \left(4\pi \frac{\omega^2}{c^2} \right) \frac{1}{c} d\omega = \frac{V\omega^2}{\pi^2 c^3} d\omega$$

- (c) Express the total internal energy U for this radiation inside the cavity as an integral over ω . Discuss analytically for very high and very low ω how the integral depends on ω . Sketch the integrand as a function of ω for two different values of T .

Using the equipartition of energy, I have classically

$$E = \int_0^{\infty} kT \left(\frac{V\omega^2}{\pi^2 c^3} \right) d\omega$$

This integral clearly diverges as ω becomes very large, and scales linearly in T .

- (d) Show that $U = aVT^4$ and give an integral expression for the constant a .

$$\langle E \rangle = \frac{\sum_i E_i e^{-\beta E_i}}{Z} = \frac{-\sum_i \frac{\partial}{\partial \beta} e^{-\beta E_i}}{Z} = -\frac{1}{Z} \frac{\partial}{\partial \beta} Z$$

$$Z = \frac{1}{1 - e^{-\hbar\omega\beta}} \quad \frac{\partial Z}{\partial \beta} = -\frac{\hbar\omega e^{-\hbar\omega\beta}}{(1 - e^{-\hbar\omega\beta})^2}$$

$$-\frac{1}{Z} \frac{\partial Z}{\partial \beta} = \frac{\hbar\omega e^{-\hbar\omega\beta}}{1 - e^{-\hbar\omega\beta}} = \frac{\hbar\omega}{e^{\hbar\omega\beta} - 1}$$

$$U = \int_0^{\infty} \frac{\hbar\omega}{e^{\hbar\omega\beta} - 1} \left(\frac{V\omega^2}{\pi^2 c^3} \right) d\omega$$

$$x = \hbar\omega\beta \quad dx = \hbar\beta d\omega$$

$$U = \int_0^{\infty} \frac{\hbar}{e^x - 1} \left(\frac{V}{\pi^2 c^3} \right) \left(\frac{x}{\hbar\beta} \right)^3 \frac{1}{\hbar\beta} dx = \frac{V}{\pi^2 c^3 \hbar^3 \beta^4} \int_0^{\infty} \frac{x^3}{e^x - 1} dx = \left[\frac{k_B^4}{\pi^2 c^3 \hbar^3} \int_0^{\infty} \frac{x^3}{e^x - 1} dx \right] VT^4$$

- (e) Use part (d) to compute the entropy S and pressure P as functions of V and T . Determine any constants that appear in terms of the symbols introduced above. Show that your results are thermodynamically consistent with the formula in part (d).

With $U = TS - PV$, $dU = TdS$ at constant V .

$$\text{Let } a = \frac{k_B^4}{\pi^2 c^3 \hbar^3} \int_0^\infty \frac{x^3}{e^x - 1} dx$$

$$\frac{dU}{T} = dS$$

$$\frac{4aVT^3 dT}{T} = dS$$

$$\frac{4}{3} aVT^3 = S$$

$$\text{With } U = TS - PV, P = -\left(\frac{\partial U}{\partial V}\right)_S$$

$$U = aVT^4 = a\left(\frac{3S}{4a}\right)^{\frac{4}{3}} V^{\frac{1}{3}}$$

$$P = -\left(\frac{\partial U}{\partial V}\right)_S = \frac{1}{3} a\left(\frac{3S}{4V}\right)^{\frac{4}{3}} = \frac{1}{3} aT^4$$

And from these results, clearly $U = TS - PV$ holds.